# ASTRONOMY 16-SUMMER 2020 

## Exercise Sheet 1

DUE by Tuesday, July 7, 2020
300 points

## A. Root Finding

A completely ionized, homogeneous hydrogen plasma is irradiated by X-ray and the electron scattering optical depth is very large. The electrons at temperature $T_{\mathrm{e}}$ are exchanging energy with protons at temperature $T_{\mathrm{p}}$ via Coulomb collisions and with photons at temperature $T_{\gamma}$ via Compton scattering. The magnetic field and all other radiative processes are negligible. Under these conditions the energy balance equation for the electron plasma can be written as:

$$
\begin{equation*}
f\left(T_{\mathrm{e}}\right)=Q \frac{U}{\rho c^{2}}\left(T_{\mathrm{e}}-T_{\gamma}\right)-\frac{\Lambda}{c^{3}}\left(\frac{T_{\mathrm{p}}}{T_{\mathrm{e}}}-1\right) \frac{1}{\sqrt{T_{\mathrm{e}}}}=0, \tag{1}
\end{equation*}
$$

where $Q \equiv 4 k_{\text {es }} K_{\mathrm{B}} / m_{\mathrm{e}} c^{2}=2.7 \times 10^{-10} \mathrm{~cm}^{2} \mathrm{~g}^{-1} \mathrm{~K}^{-1}$ (with $k_{\mathrm{es}}=0.4 \mathrm{~cm}^{2} \mathrm{~g}^{-1}$ being the electron scattering opacity), $U /\left(\rho c^{2}\right)$ is the ratio between the radiation energy density and the rest-mass energy density of the electron-proton plasma, and $\Lambda=4.4 \times 10^{30}$ is a constant.
(i) Consider the case in which the electrons are heated by Coulomb collisions with the protons at temperature $T_{\mathrm{p}}=10^{9} \mathrm{~K}$ and cooled via inverse Compton scattering with photons at temperature $T_{\gamma}=10^{7} \mathrm{~K}$. Solve Eq. 1 with $U /\left(\rho c^{2}\right)=1$ and find the equilibrium temperature for the electrons $T_{\mathrm{e}}$. The bracketing interval can be found after plotting Eq. 1 as a function of $T_{\mathrm{e}}$.
(ii) Consider now the case in which the electrons are cooled by Coulomb collisions with the protons at temperature $T_{\mathrm{p}}=10^{7} \mathrm{~K}$ and heated via Compton scattering with photons at temperature $T_{\gamma}=10^{9} \mathrm{~K}$. Solve Eq. 1 with $U /\left(\rho c^{2}\right)=8 \times 10^{-5}$ and find the equilibrium temperature for the electrons using as bracketing interval $T_{1}=10^{7} \mathrm{~K}$ and $T_{2}=10^{9} \mathrm{~K}$

NOTE:
(i) Use both the bisection and the Newton-Raphson methods concentrating on the number of iterations necessary to reach the desired accuracy of $\epsilon=10^{-7}$.
(ii) To make results comparable, use as speed of light $c=2.99 \times 10^{10} \mathrm{~cm} \mathrm{~s}^{-1}$

## B. Polynomial Interpolation

Consider the function

$$
\begin{equation*}
y(x)=3+200 x-30 x^{2}+4 x^{3}-x^{4} \tag{2}
\end{equation*}
$$

in the interval $I: x \in[-10,10]$.
(i) Find its roots in $I$ if they exists (check whether roots exist by plotting Eq. 2). Use, as bracketing intervals, e.g., $y_{1}=-1$ and $y_{2}=1$ for one root, and $y_{3}=4$ and $y_{4}=6$ for the other root.
(ii-BONUS) Interpolate the value of the function at $x=-5$ and $x=5$ using a linear, a quadratic, and a cubic interpolating polynomial. Calculate the error made in each case. Write the adopted input values, the results, and the errors in tables for each interpolating methods. For the linear polynomial, use $x_{1}=-6$ and $x_{2}=-4$ for $x=-5$, and $x_{1}=4$ and $x_{2}=6$ for $x=5$. For the quadratic polynomial, use $x_{1}=-6, x_{2}=-4$, and $x_{3}=-3$ for $x=-5$, and $x_{1}=4, x_{2}=6$, and $x_{3}=7$ for $x=5$. For the cubic polynomial, use $x_{1}=-7, x_{2}=-6, x_{3}=-4$, and $x_{4}=-3$ for $x=-5$, and $x_{1}=3, x_{2}=4, x_{3}=6$, and $x_{4}=7$ for $x=5$.

