AST31 - Stellar Structure and Evolution

AST-31 is a course intended for students with a background in physics and/or astronomy and mathematics and a strong interest in astronomy. The course will cover the following topics: interaction of light and matter, telescopes, stellar structure, stellar evolution, interstellar medium, exoplanets, and formation of planetary systems. The course will also include assignments involving the use of computers.

May be used to satisfy the concentration requirements for major and minors in the Department of Physics and Astronomy.

Prerequisites: PHY11/12 or consent of instructor.





WHEN: Fall 2020, Block T+: TueThu, 9:00-10:15am WHERE: Zoom Contact: Prof. Danilo Marchesini

For more information: email: <u>danilo.marchesini@tufts.edu</u> <u>http://cosmos.phy.tufts.edu/~danilo/AST31/AST31.html</u>



- 1. Brief description of the Astronomy group/research @ Tufts
- 2. Overview of the course, syllabus, schedule
- 3. Survey (15 min)
- 4. Q&A

Interested in doing research in Astronomy?

Contact:



Prof. Anna Sajina (Anna.Sajina@tufts.edu) <u>http://cosmos2.phy.tufts.edu/~asajina/</u>

Prof. Danilo Marchesini (Danilo.Marchesini@tufts.edu) http://cosmos.phy.tufts.edu/~danilo/



Astronomy Group

GRADUATE STUDENTS:

Cemile Marsan — defended Aug. 2017 —> now postdoc in Toronto Andrea Silva-Bustamante — defended Jan. 2018 —> now postdoc in Tokyo Nick Martis — defended Apr. 2020 —> now postdoc in Victoria (Canada)

Kalina Nedkova (6th-yr) Kumail Zaidi (3rd-yr) Max Kurzner (2nd-yr) Mudit Garodia (2nd-yr) Richard Pan (1st-yr) Valentina La Torre (1st-yr)

UNDERGRADUATE STUDENTS (currently involved in research):

Lilianna Houston Teddy Peña Victor Soto-Castro Jacob Whitney How did the tiny fluctuations in the distribution of matter we see in the Cosmic Microwave Background Radiation grew into galaxies?

Today

380,000 years after the Big Bang

Co-evolution of SMBH ? and host galaxy



Prof. Marchesini

Understanding how galaxies formed shortly after the Big Bang, and how they evolved with cosmic time.

To do this I search for and study the most distant galaxies in the universe...



Hunting monsters: searching for Very Massive Galaxies when the Universe was younger than 2 billion years old (Marchesini+2010)

Image credits: Danilo Marchesini (Tufts University)

KECK TELESCOPES: HAWAII



VERY LARGE TELESCOPES: CHILE

MAGELLAN TELESCOPE: CHILE



GRANTECAN: CANARY ISL., SPAIN



Confirmation of the Existence of Very Massive Galaxies in the Early Universe (<u>Marsan+2015/2017</u>)



Stellar mass ~ 300 billion M_{Sun} (4x more massive than MW) when the Universe is only ~1.8 billion years old.

Age of stars in this galaxy ~ 400 million years

Most of the stars formed in a very short and intense burst of star formation

Collaboration expanded to include University of California - Riverside/Irvine + York University (Toronto, CA); tens of Keck nights awarded (Forrest+2020a/b)

The Hubble Frontier Fields program







Catalogs published in <u>Shipley et al. (2018)</u> and used by Kalina+Victor for their projects

Prime Focus Spectrograph (PFS)





Prime Focus Spectrograph



PFS Website: https://pfs.ipmu.jp/intro.html



Professor Anna Sajina Astrophysics

My research involves studies of distant, dusty IR-bright galaxies. Galaxies like that often also host active galactic nuclei which are growing supermassive black holes. This research gives insight into how galaxies formed and evolved from the Big Bang to the present and what is the role of the supermassive black holes in the centers of galaxies.



Here are the tools I use in my research.....

Infrared telescopes like the Spitzer Space Telescope



Radio telescopes, like the VLA in New Mexico



Simulated observations based on N-body galaxy simulations



And couldn't do without my trusty laptop — astrophysics research involves plenty of coding!



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ASSIGNMENT for Lecture 2 (Thursday, September 10):

Read and understand Ch. 2 (2.1, 2.2, and 2.3; skip 2.4) - this may take 2-3 hours if done properly, although you should have seen all of this before (PHY11). The demonstration in Example 2.2.1 is facultative. Try to understand all other demonstrations, although we will also discuss them during the lectures.